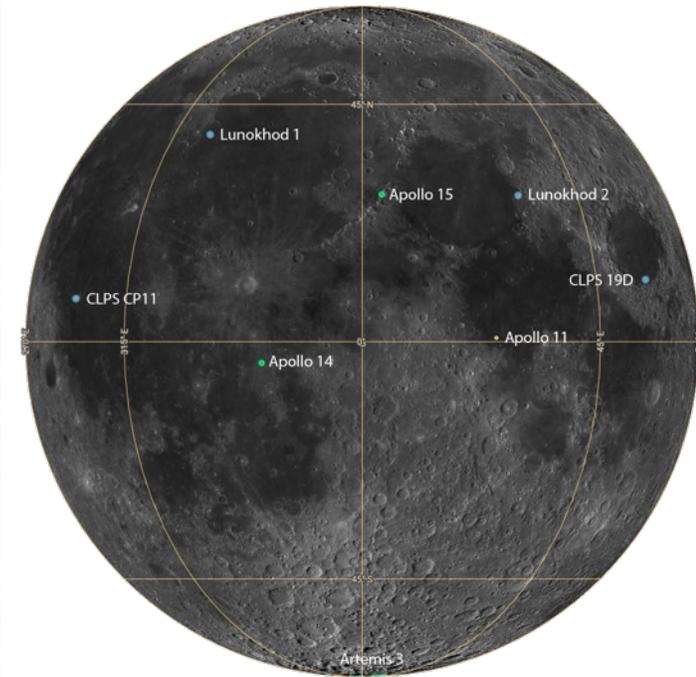
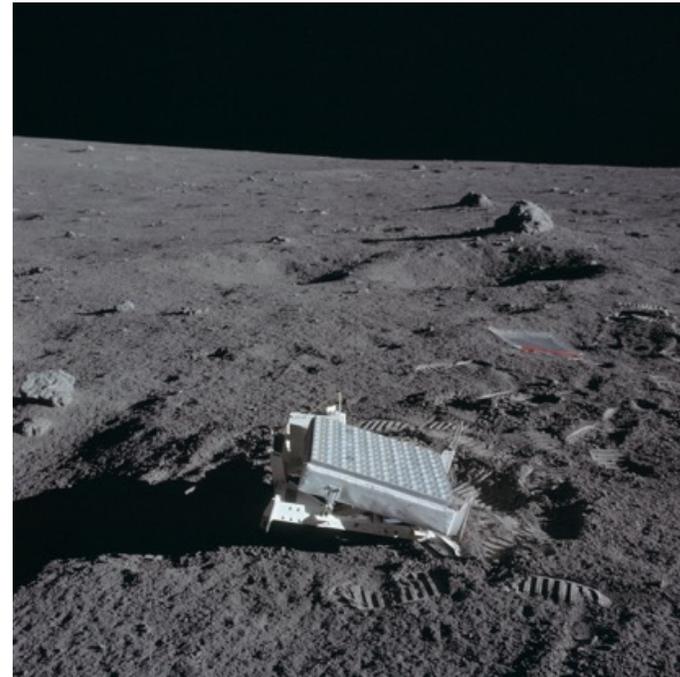




International Lunar Reference System Next Steps

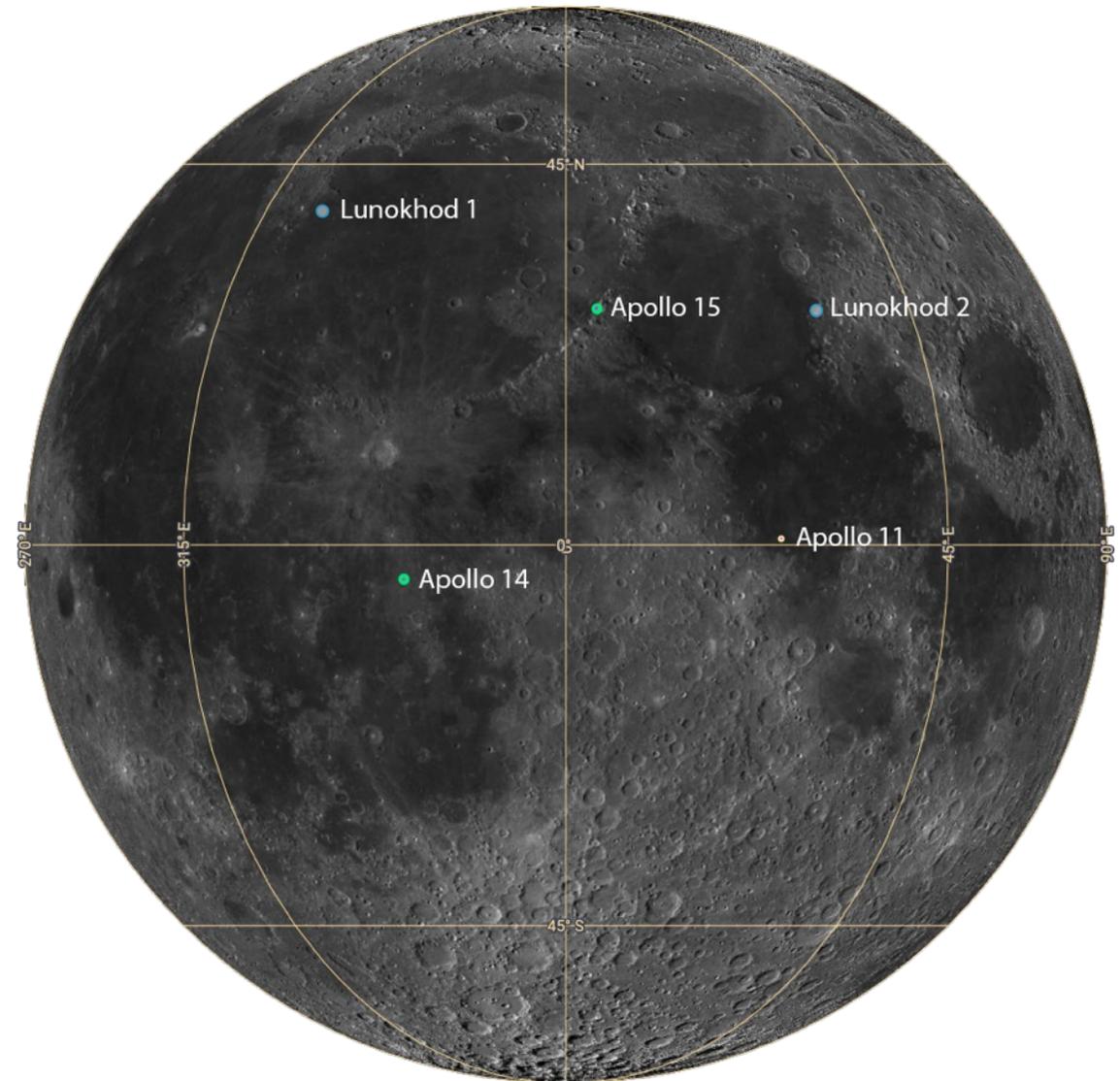
Stephen Merkowitz
NASA/GSFC

2nd Joint ICG-IOAG Multilateral Cislunar PNT Workshop
February 11, 2026



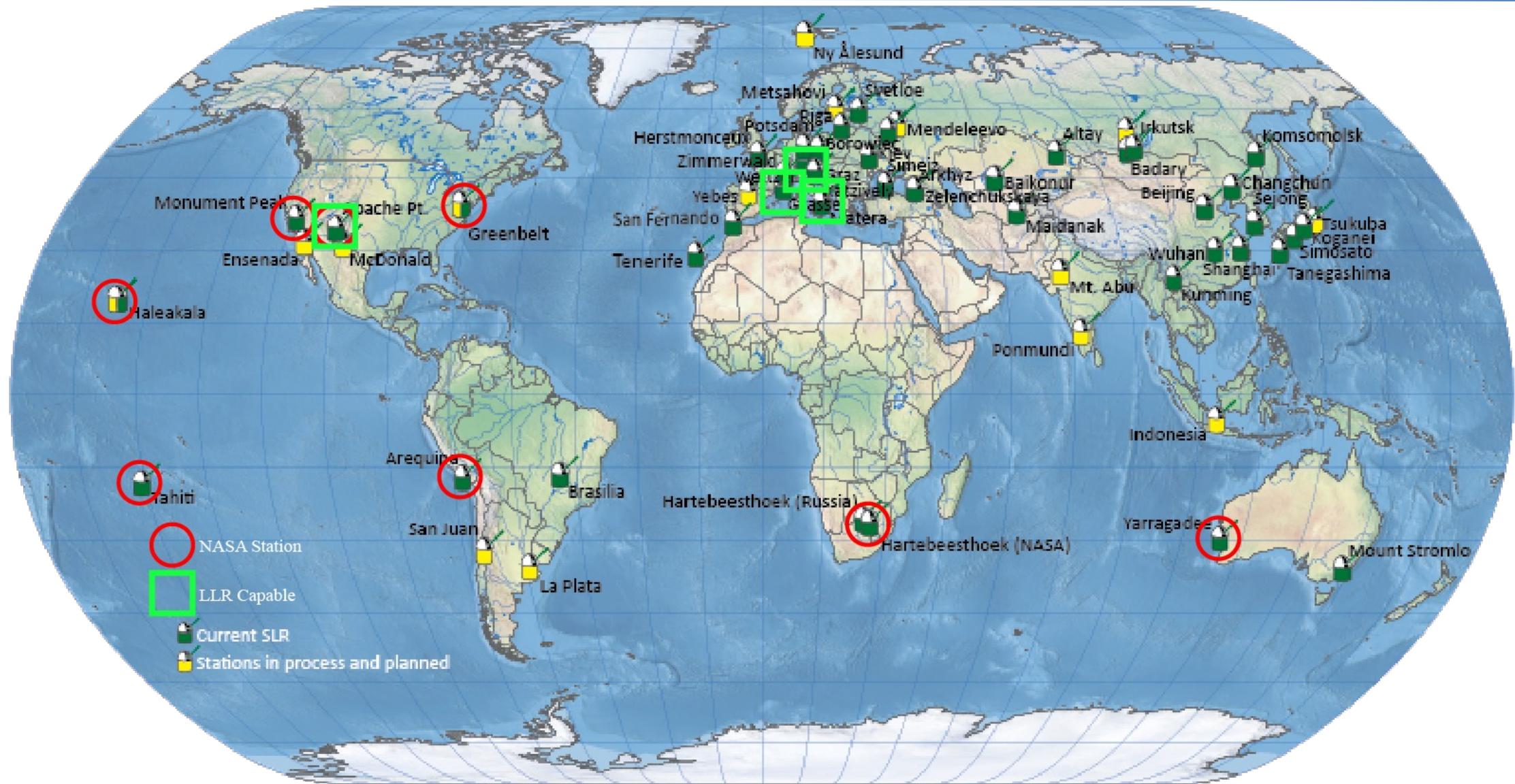
Current Lunar Retroreflectors

- ◆ All the lunar retroreflectors used in the current LRS lie within 26 degrees latitude of the equator, and the most useful ones (Apollo) within 24 degrees longitude of the sub-earth meridian. This clustering weakens their geometrical strength.
- ◆ The Apollo and Lunokhod retroreflectors continue to be useful targets, but are showing signs of degradation, most likely due to dust.



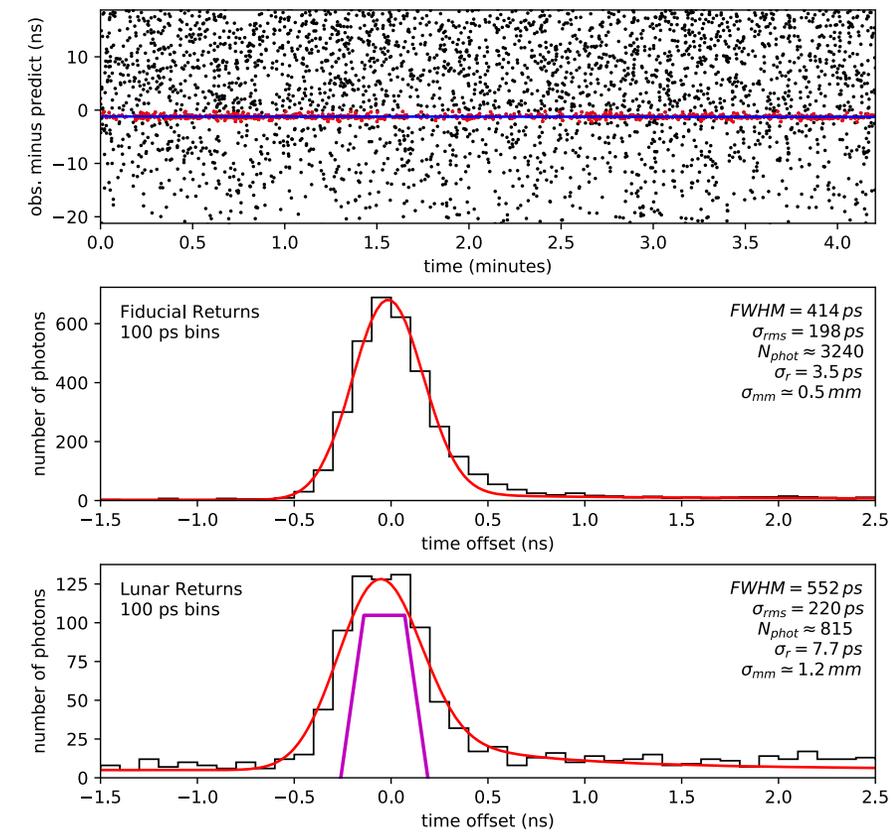
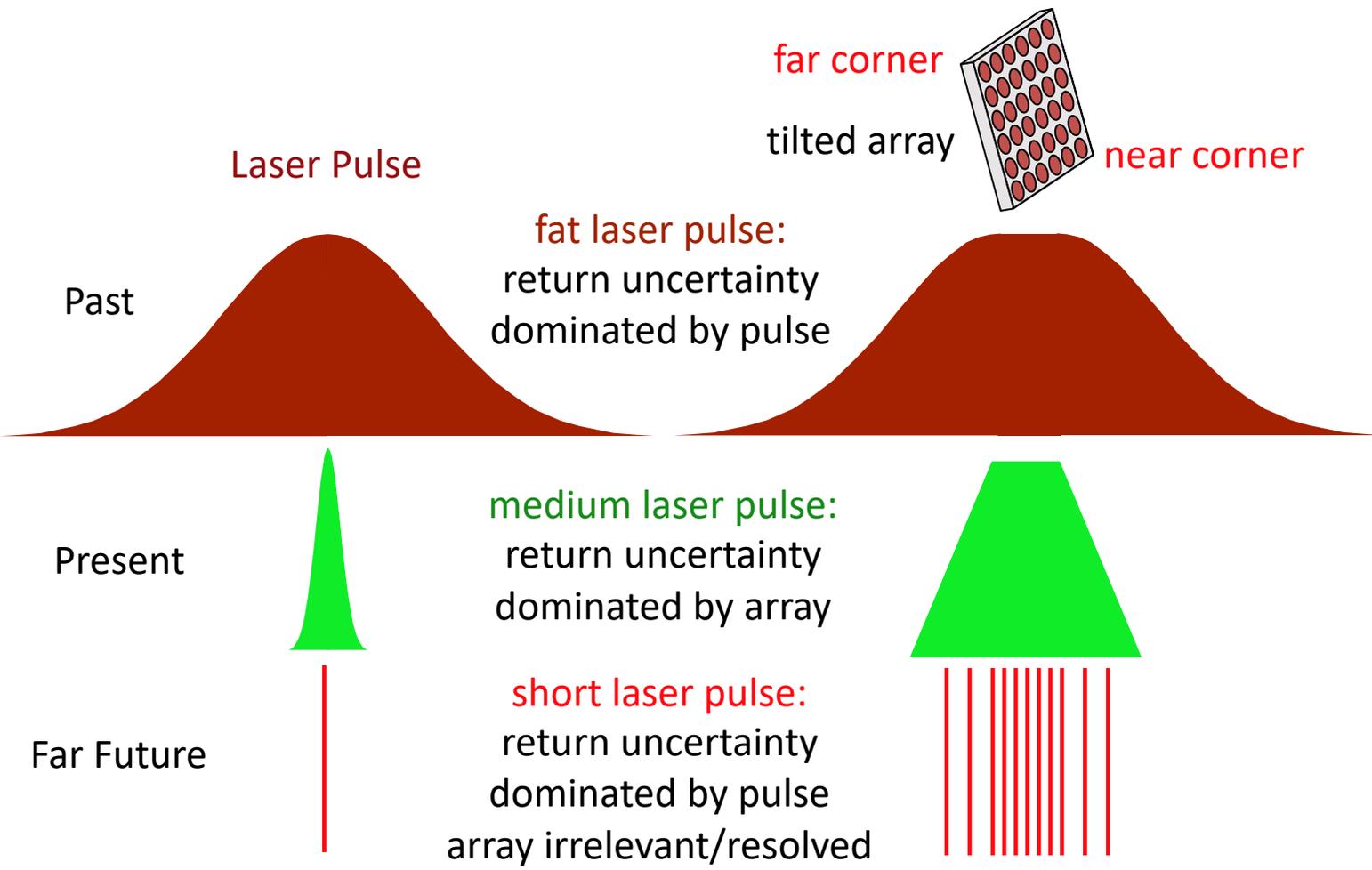


International Laser Ranging Service (ILRS) Stations





Current Limiting Error - Retroreflector Array Tilt



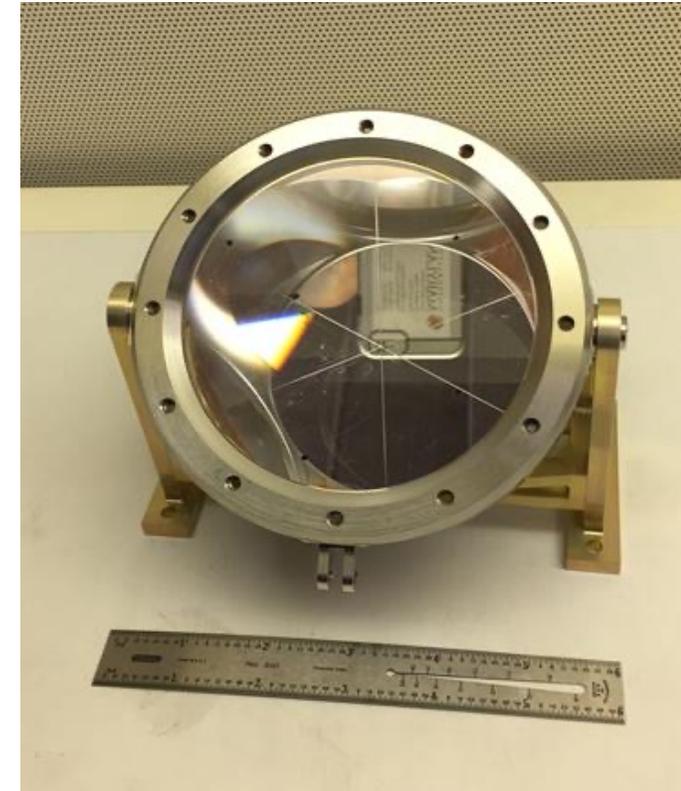
Apache Point measurements with retroreflector array timing profile

Colmenares et al.
<https://doi.org/10.1088/1538-3873/acf787>



Solution to Tilt Error – Single Cubes

- ◆ Single cubes do not broaden return pulse thus eliminating the array tilt error.
- ◆ Central irradiance of return goes up as fourth power of the cube corner diameter, making it feasible to get comparable optical cross section to an Apollo 100 x 3.8 cm cube array with a single 12 cm cube.
- ◆ Solution does not come for free:
 - Solid cubes are susceptible to thermal and polarization effects that can impact their performance.
 - Tuning the cubes for the lunar motion (velocity aberrations) is also a challenge for single cubes versus an array.

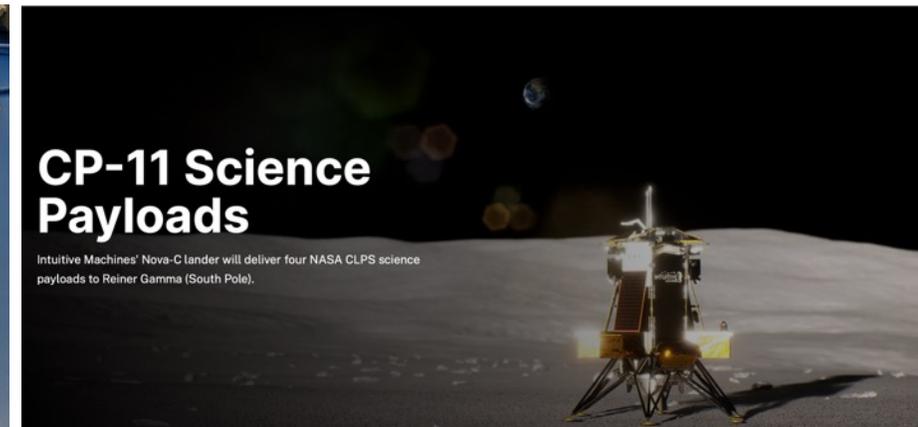
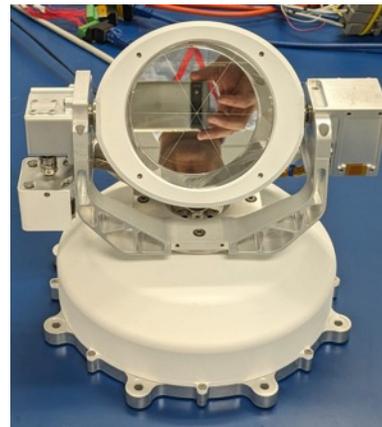
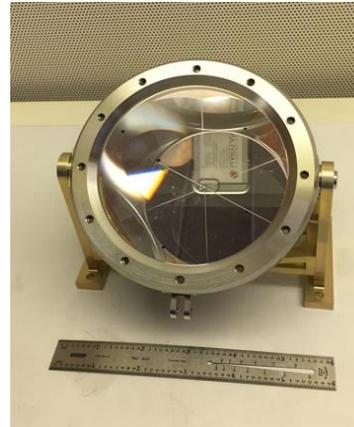


NGLR courtesy Douglas Currie



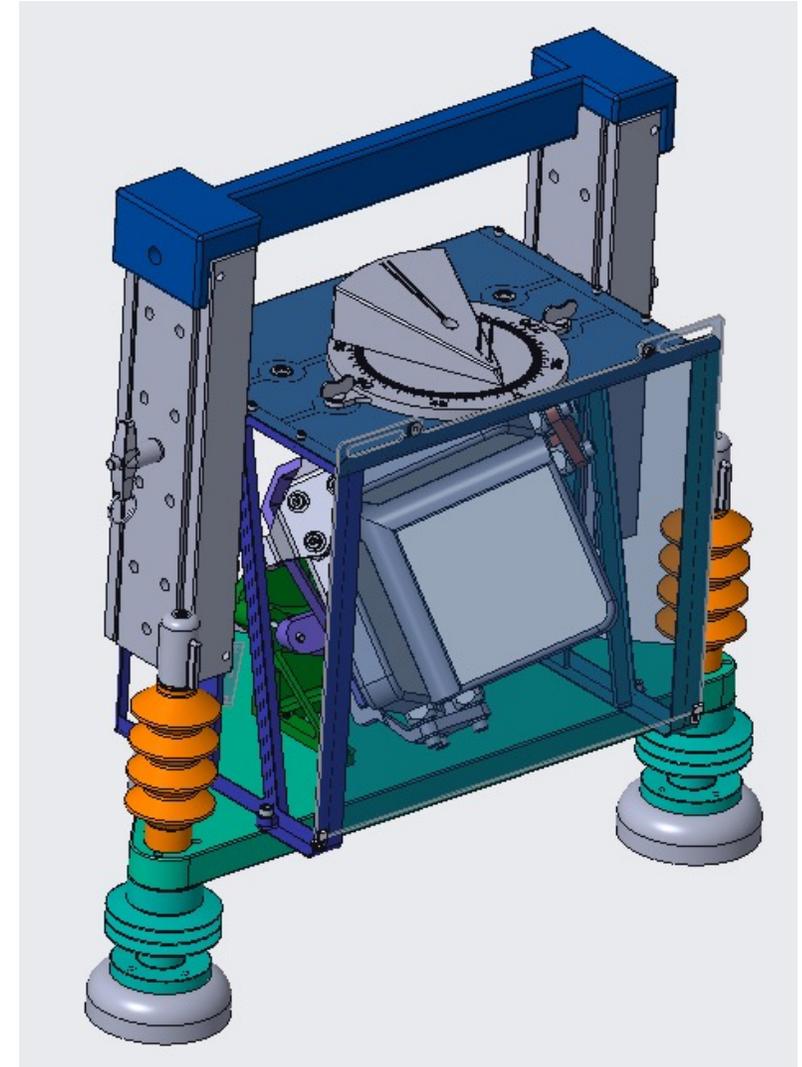
NASA Commercial Lunar Payload Services (CLPS) Deliveries

- ◆ Firefly's Blue Ghost 1 lander to Mare Crisium
 - University of Maryland's Next Generation Lunar Retroreflector (NGLR), Dr. Douglas Currie, Principal Investigator
 - Successfully landed March 2, 2025. LLR data from NASA's Apache Point station available from NASA's CDDIS.
- ◆ Intuitive Machines' Nova-C lander to Reiner Gamma
 - ESA's MoonLIGHT Pointing Actuator (MPAc)
- ◆ Both CLPS missions use a single 10cm solid cube corner.



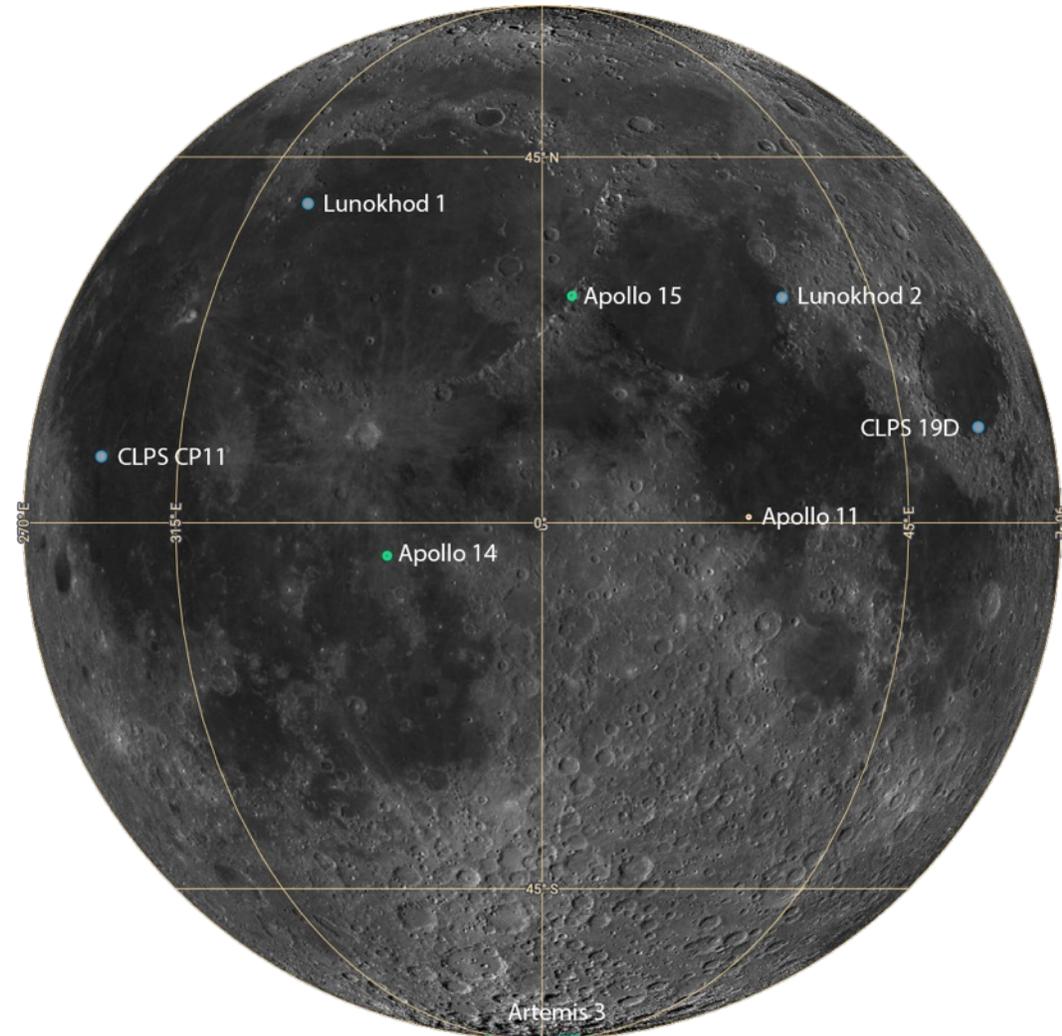
Artemis Lunar Laser Retroreflector (ALLR)

- ◆ NASA Goddard Space Flight Center developed a single large hollow retroreflector for NASA's Artemis 3 mission.
- ◆ The Artemis 3 astronauts will deploy ALLR at the lunar south pole.
- ◆ ALLR is expected to have a significantly greater optical cross section than the Apollo 11 array.



Upcoming LLR Distribution

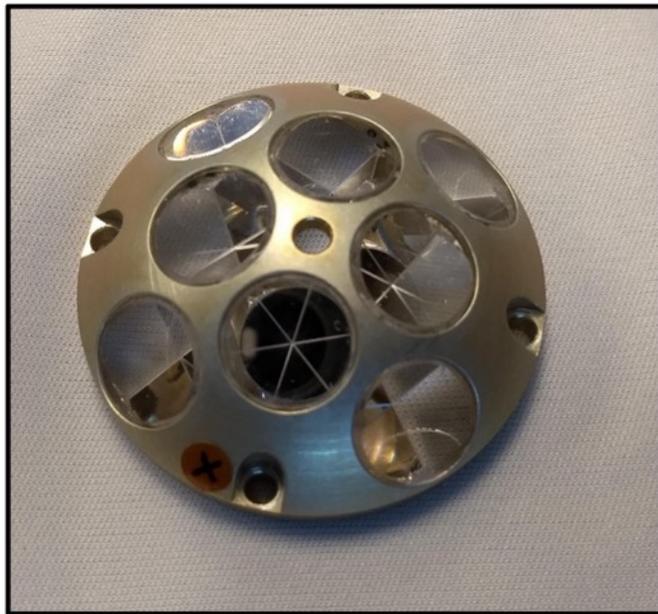
- ◆ Next-generation lunar retroreflectors deployed to new lunar sites on the NASA CLPS and Artemis missions greatly expand the geometric coverage and build upon the long legacy of LLR contributions to the lunar reference system.



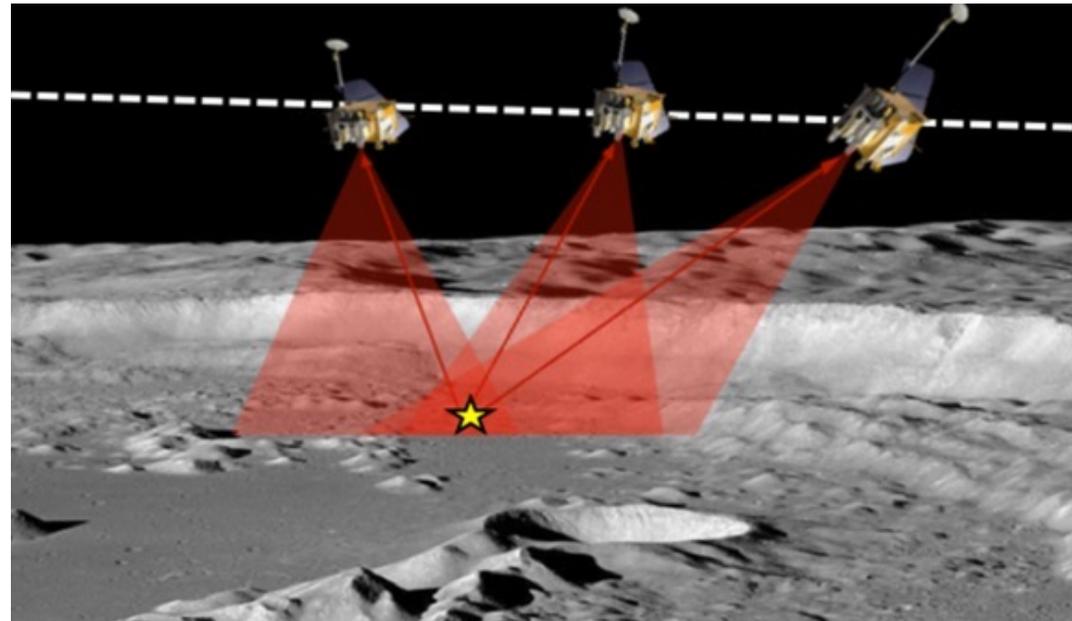


Miniature Laser Retroreflector Array (mini-LRA)

- ◆ NASA Goddard Space Flight Center is also putting miniature retroreflectors on most CLPS mission that can be seen by lunar altimeters.
- ◆ The Lunar Reconnaissance Orbiter (LRO) Lunar Orbiter Laser Altimeter (LOLA) has successfully.
- ◆ Mini-LRA's provide potential additional reference points for future reference frame realizations.



5.1x1.6 cm,



Xiaoli Sun et al. (2019), <https://doi.org/10.1364/ao.58.009259>



Lunar Reconnaissance Orbiter (LRO) Lunar Orbiter Laser Altimeter (LOLA) Retroreflector Dataset

as of Jan.1, 2026

The larger LRAs (Apollo, Lunokhod) have larger optical cross-sections which increases the chance of return (50-100%).

The smaller arrays have more variation, also due to geometry, but generally in line or higher than predicted (~20%).

2023-2024 dataset available:
[doi:10.60903/gsfcpgda-lola-lra](https://doi.org/10.60903/gsfcpgda-lola-lra)
 (expect update with 2025 data)

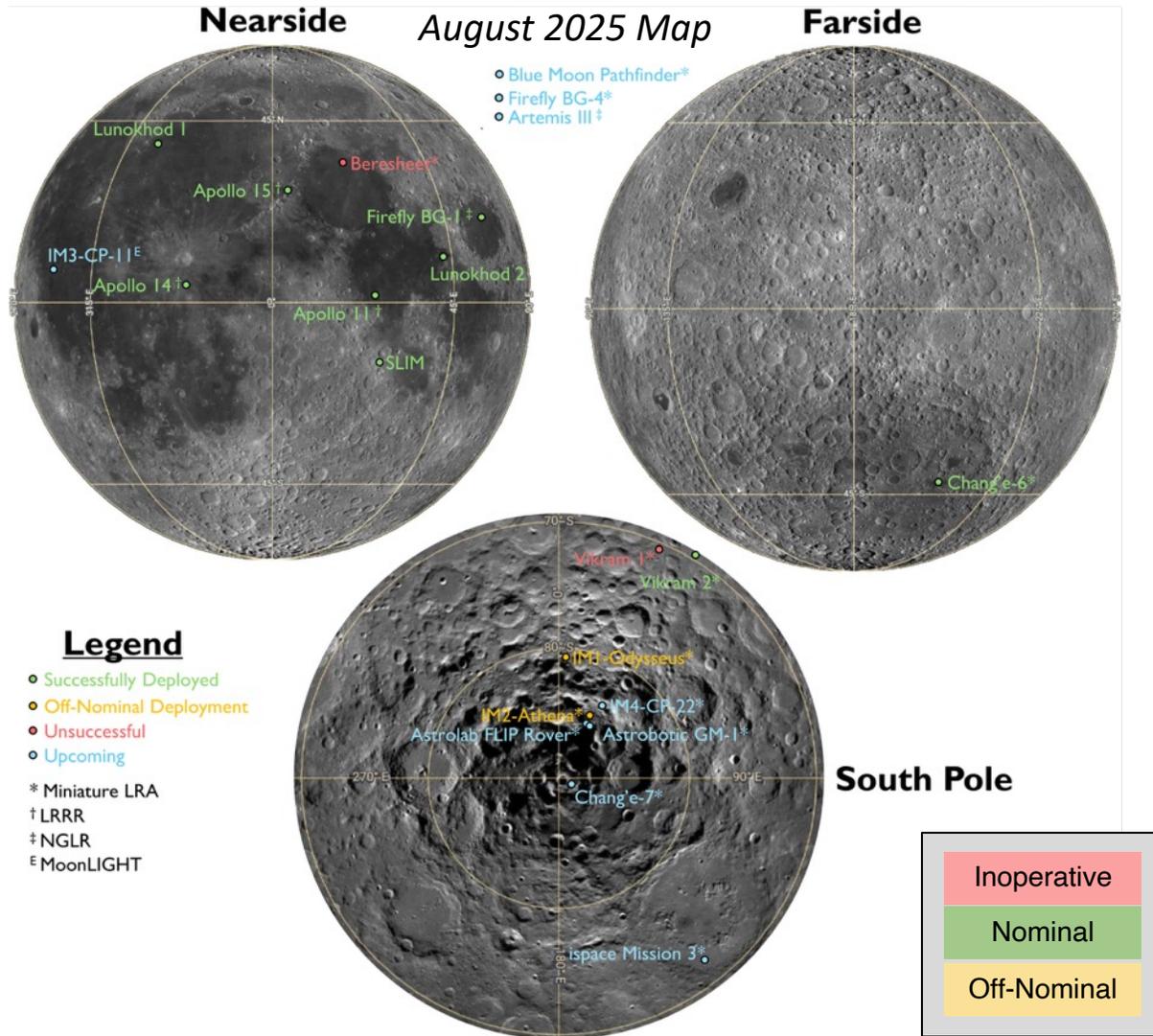


Array	Total Attempts	Total Successes	Success Rate
Apollo 11	13	10	77%
Apollo 14	30	26	87%
Apollo 15	15	8	53%
Chandrayaan-3	135	25	19%
SLIM	15	4	27%
Chang'E-6	59	12	20%
BG-1	5	4	80%
Lunokhod-1	6	5	83%
Lunokhod-2	4	4	100%
IM-1	18	0	0%

Credit: Erwan Mazarico (GSFC) & LOLA Science Team



Mini-LRA Distribution



Mission (Org.)	Launch Date	Surface Location
Beresheet (Spacell)	February 21, 2019	32.5956°N, 19.3496°E
Chandrayaan-2 (ISRO)	July 22, 2019	70.8810°S, 22.7849°E
Chandrayaan-3 (ISRO)	July 14, 2023	69.3735°S, 32.3199°E
SLIM (JAXA)	September 6, 2023	13.31549°S, 25.24889°E
CLPS TO-2AB (Astrobotic)	January 8, 2024	N/A
CLPS TO-2IM (Intuitive Machines-1)	February 15, 2024	80.1277°S, 1.4357°E
<i>Chang'e-6 (CNSA/ESA)</i>	May 3, 2024	41.6385°S, 153.9852°E
CPLS PRIME-1 (Intuitive Machines-2)	February 26, 2025	84.7754°S, 29.1264°E
CLPS TO-20A GM-1 (Astrobotic)	2026	Mons Mouton
FLIP Rover (Venturi Astrolab)	2026	Mons Mouton
Blue Moon Mk-1/CT-3 (Blue Origin)	2026	South Pole Region
<i>Chang'e-7 (CNSA/ESA)</i>	2026	Shackleton Rim
CLPS CP-11 (Intuitive Machines-3)	2026	Reiner Gamma
CLPS CP-22 (Intuitive Machines-4)	2027	Mons Mouton
<i>ispace Mission 3</i>	2027	Schrödinger Basin
Artemis III (NASA)	2028	South Pole Region
CLPS CS-6 (Firefly BGM4)	2029	South Pole Region

Updates of this map and list for the NASA GSFC mini-LRAs are available: [doi:10.60903/GSFC-miniLRAs](https://doi.org/10.60903/GSFC-miniLRAs)

VLBI to Lunar Landers

- ◆ Radio beacons on the lunar surface can be observed by VLBI along with nearby quasars to provide a transverse component. LLR provides the radial component.
- ◆ Around 1W power (similar to cell phone) is sufficient for observations.
- ◆ Expected accuracies of $\sim 3\text{cm}$ with 5 hours of observations and $\sim 1\text{cm}$ after 1 month.
- ◆ VLBI is not susceptible to weather like LLR.

